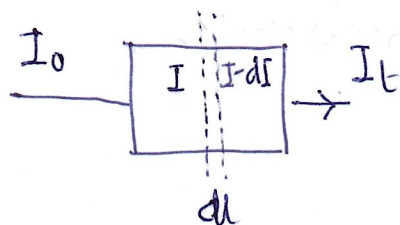


Molecular Spectroscopy

Absorption spectroscopy: Absorbance $A = \log \frac{I_0}{I_t}$

Emission spectroscopy: Emission quantum yield $= \frac{I_{em}}{I_{abs}} = \phi_{em}$



$$-dI \propto I \cdot c \cdot dl$$

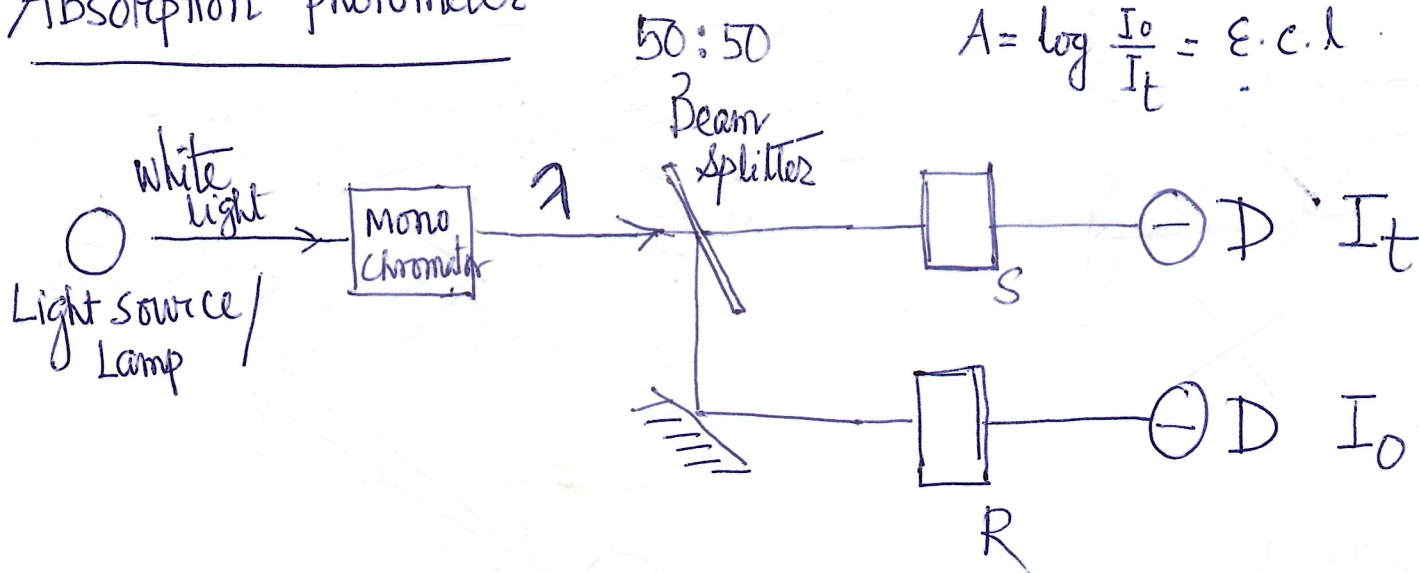
$$-dI = a \cdot I \cdot c \cdot dl$$

$$-\frac{dI}{I} = a \cdot c \cdot dl \Rightarrow \int_{I_t}^{I_0} \frac{dI}{I} = a \cdot c \cdot \int_0^l dl$$

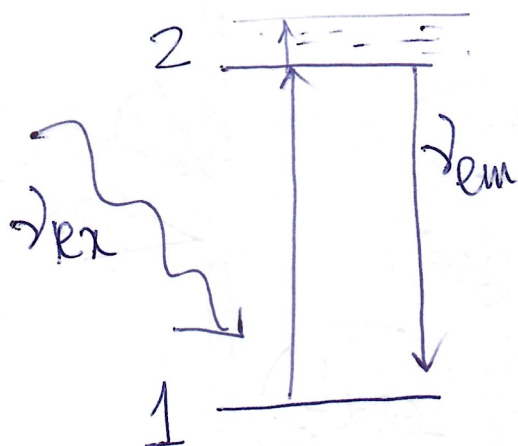
$$\ln \frac{I_0}{I_t} = a \cdot c \cdot l$$

$$A = \log \frac{I_0}{I_t} = \epsilon \cdot c \cdot l$$

Absorption photometer

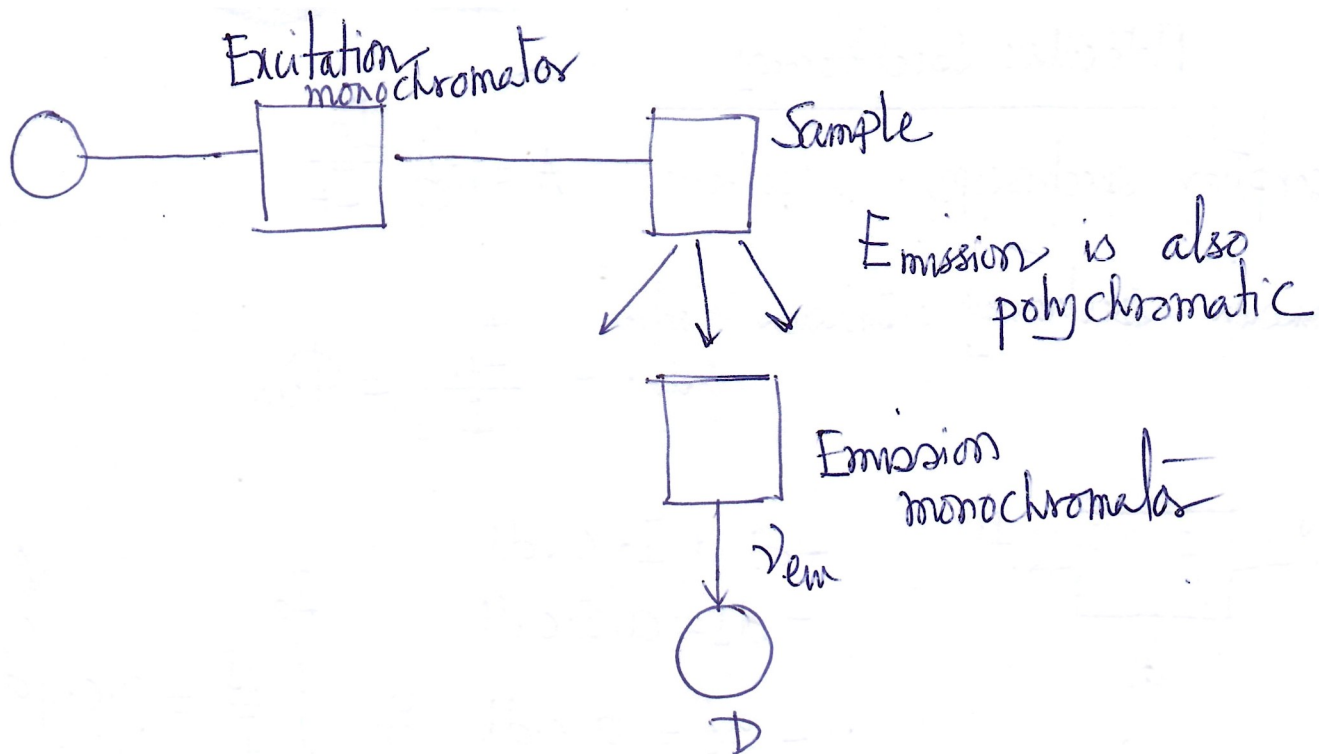


Emission photometer

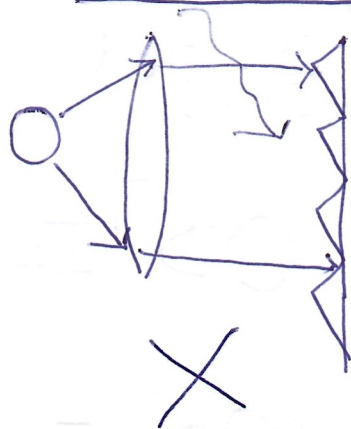


Generally λ_{ex} & λ_{em} are different
 λ_{ex} should be chosen in such a way that the sample is excited from state 1 to state 2.

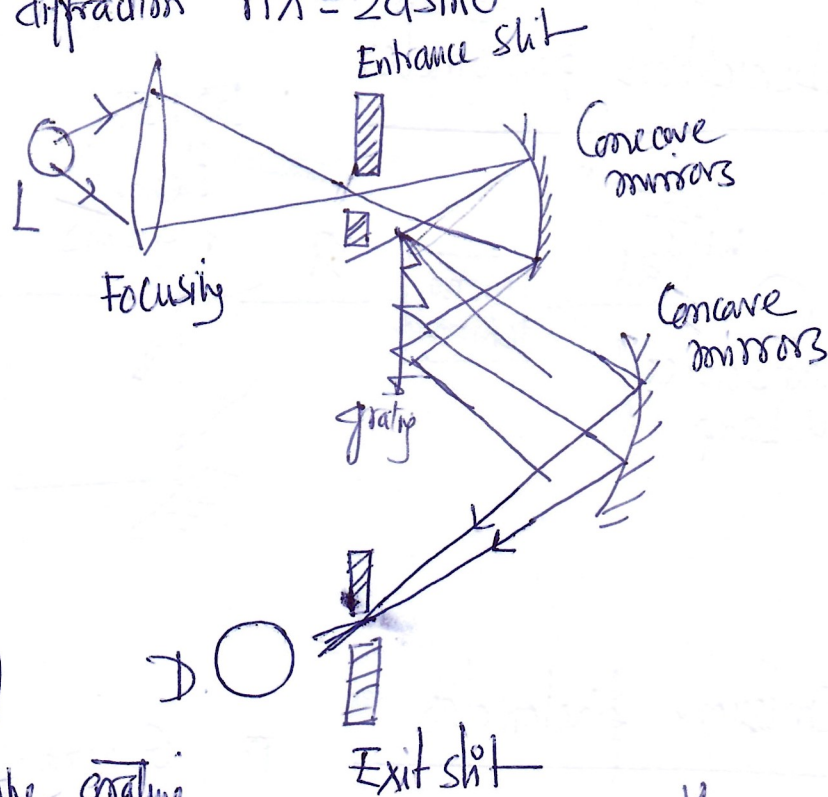
Generally, energy which is emitted is lower than the excitation.



Monochromator



Bragg's diffraction $n\lambda = 2d\sin\theta$



Smaller d
better resolution

We can only tilt the angle of the grating



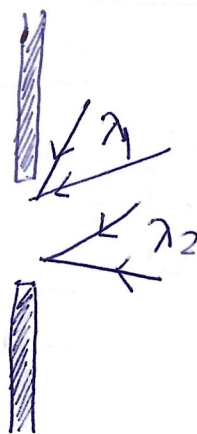
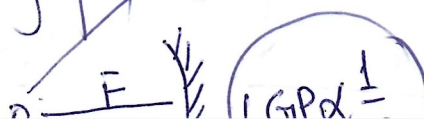
Too good is no good

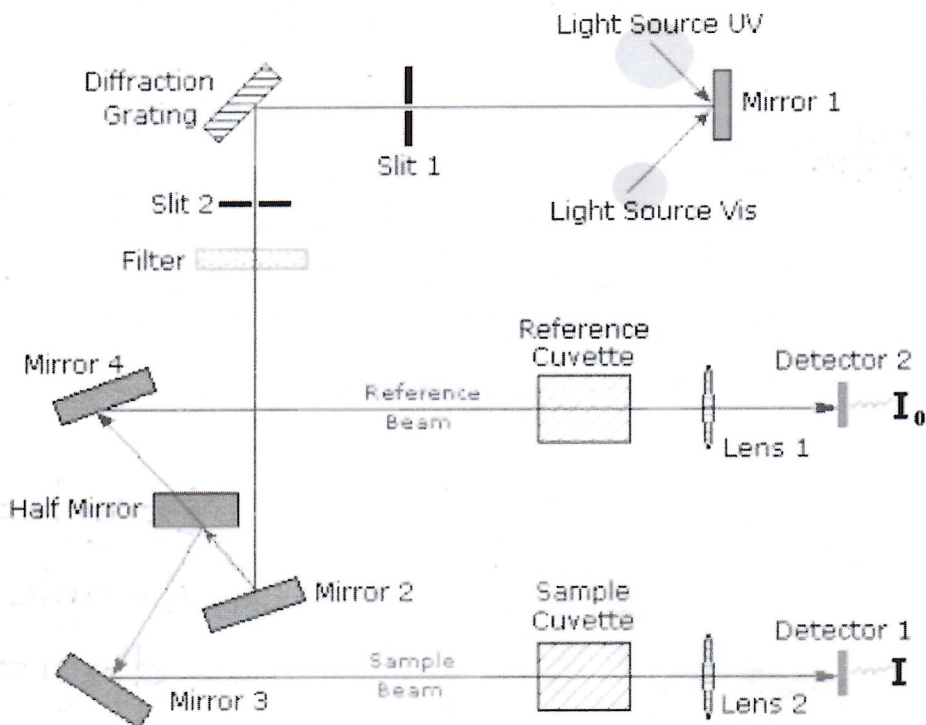
$$\lambda_1 = 2d\sin\theta_1$$

$$\lambda_2 = 2d\sin\theta_2$$

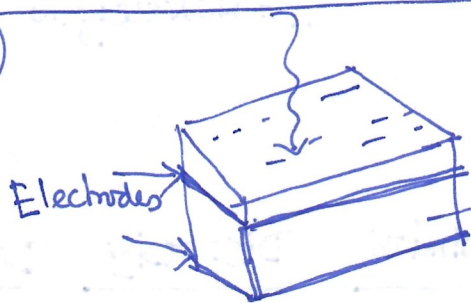
$$\lambda_2 - \lambda_1 = 2d \cos\theta \Delta(\sin\theta)$$

Light gathering power (LGP)





IR



Pyroelectric material

DTGS Deuterated
Triglycerine sulphate
(DTGS)

MCT Photoconductive
(Mercury Cadmium Telluride)

Resolution

The resolution of an instrument tells us how well two close-lying energies (or wavelength or freq. or masses) can be resolved.

Resolving power

The spectral resolving power of a spectrometer is a measure of its ability to resolve features

Mirrors
 Collimators
 Filters
 Modulators
 lock-in amplifier

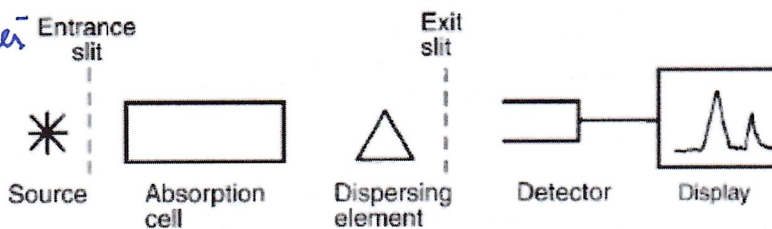
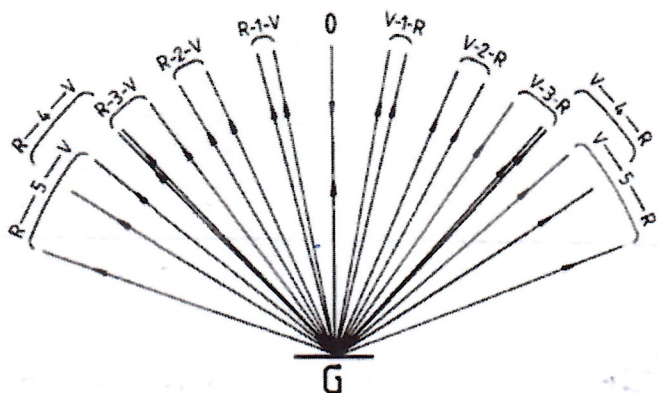


Figure 3.2 The components of a typical absorption experiment

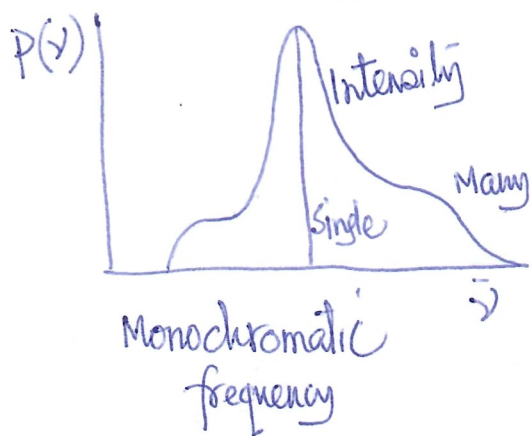
$$m\lambda = d(\sin i + \sin \theta)$$



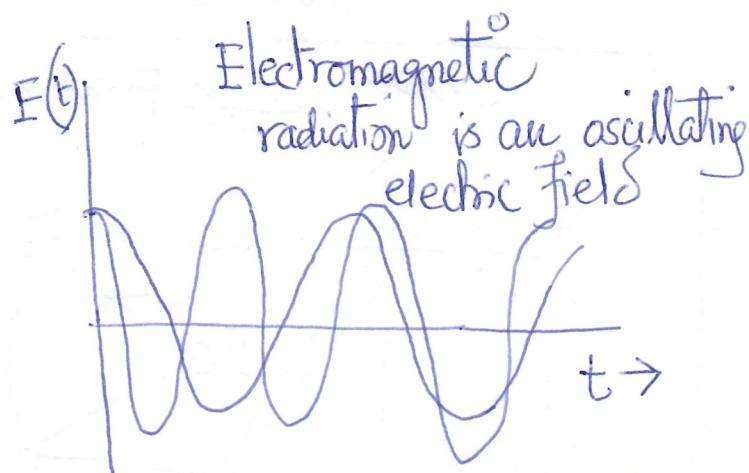
i = angle of incidence
 θ = angle of reflection
 d = groove distance
 $m = 0, 1, 2, 3, 4$
 λ = wavelength

Region	Source	Absorption Cell window	Dispersing	Detector
Microwave	Klystron	Mica	None	Crystal diode
Mid Near IR	Neonst glow filament Globars (SiC) Nichrome wire	NaCl or KBr	Grating Interferometer	Transducer DTGS MCT
Vis. UV	(vis) Tungsten-halogen (uv) Deuterium (uv-vis) Xenon lamp	Glass Quartz (uv) Plastic	Prism Grating	photo multiplier

Time domain spectroscopy



≡

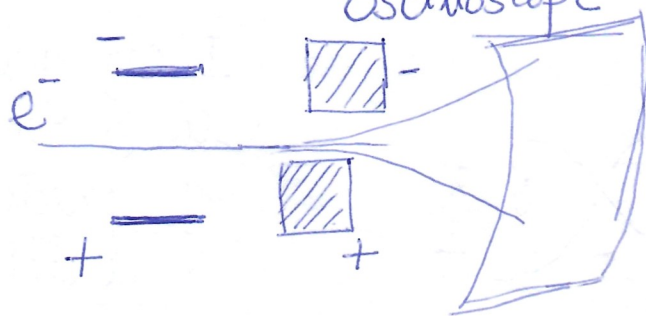


Advantages of time domain At a particular time, information of all frequencies are present.

1. Jacquinot (throughput advantage)
2. Resolution
3. Multiplex detection S/N ratio is better (all wavelengths together)

Detectors

to detect a time domain oscilloscope



$$c\bar{\nu} = \nu$$

$$1000 \text{ cm}^{-1} = \bar{\nu}$$

$$\nu = 3 \times 10^8 \times 10^2 \text{ cm}^{-1} \times 10^3 \text{ cm}^{-1} = 3 \times 10^{13} \text{ s}^{-1}$$

Wavelength
= λ

Wave number
= $\frac{1}{\lambda}$

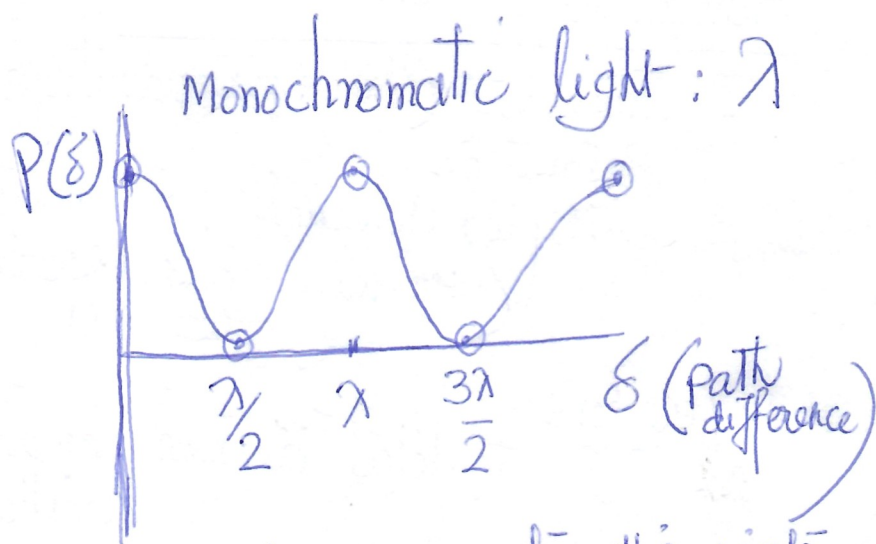
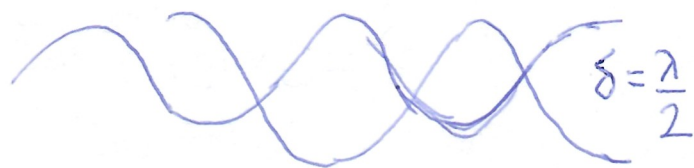
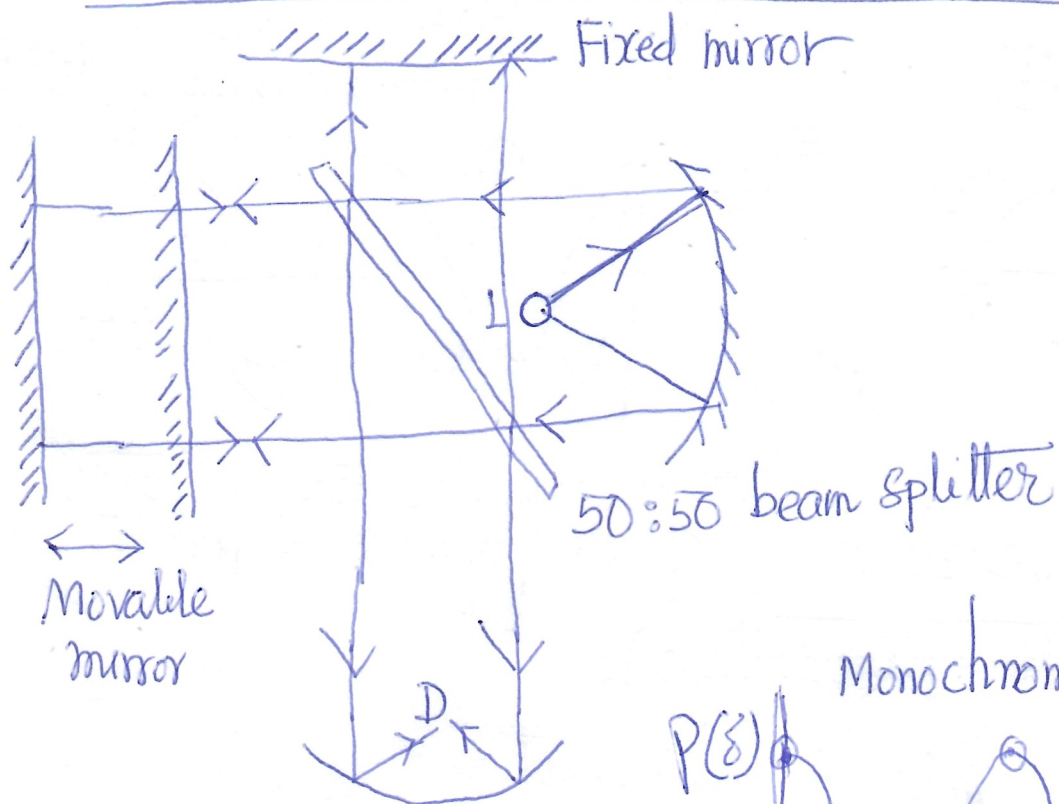
= $\bar{\nu}$

$E = h\nu$
 $= h \cdot \frac{c}{\lambda}$

Time period $\tau = \frac{1}{\nu} = 3.3 \text{ ps}$

We do not have faster oscilloscope.

Frequency modulation for Time domain



I can generate this points in δ in real measurements of practical timescale of a detector.

V_M = velocity of the mirror (M)
 τ = Time for mirror to travel $\frac{\lambda}{2}$

$$V_M \cdot \tau = \frac{\lambda}{2}$$

$$\tau = \frac{\lambda}{2V_M}$$

$$f = \frac{1}{\tau} = \frac{2V_M}{\lambda} \text{ light}$$

$$= \frac{2V_M}{c} \text{ } \downarrow$$

frequency of signal.

$$f = 2V_M \lambda^{-1}$$

How do we get resolution?

No monochromator
 No grating
 No slit

Semi-Classical Treatment of Matter-Light Interaction

Light Electromagnetic radiation

Electric field is taken into consideration

Let us consider $\Psi(x, y, z, t) = \psi(x, y, z) \phi(t) = \psi \cdot \phi$

$$\hat{H}^{(0)} = -\frac{\hbar^2}{2m} \nabla^2 + V(x, y, z)$$

Time Independent

$$\hat{H}^{(0)} \Psi = -\frac{\hbar}{i} \frac{\partial}{\partial t} \Psi$$

$$\hat{H}^{(0)} [\psi \phi] = -\frac{\hbar}{i} \frac{\partial}{\partial t} [\psi \phi]$$

$$\phi \hat{H}^{(0)} \psi = \left[\frac{\hbar}{i} \frac{\partial \phi}{\partial t} \right] \psi$$

$$[\div \psi \phi] \Rightarrow \frac{\hat{H}^{(0)} \psi}{\psi} = \left[-\frac{\hbar}{i} \frac{\partial \phi}{\partial t} \right] / \phi$$

$$\Rightarrow -\frac{\hbar}{i} \frac{\partial \phi}{\partial t} = \frac{E \psi}{\psi} \cdot \phi$$

$$\Rightarrow \frac{\partial \phi}{\partial t} = -\frac{i E \phi}{\hbar}$$

$$\boxed{\phi = e^{-i E t / \hbar}}$$

$$\text{So, } \Psi = \psi e^{-i E t / \hbar}$$

Spectroscopy means transition from one state to the other.

One photon rule: one photon brings about only one transition.

$$E_m \text{ --- } \Psi_m$$

Two level system

Superposition of mixing of states

$$E_L \text{ --- } \Psi_L$$

$$\Psi = C_L \Psi_L + C_m \Psi_m$$

Mixing coefficients C_L, C_m

At $t \rightarrow 0, C_L = 1, C_m = 0$

No light, $t=0, \hat{H} = \hat{H}^{(0)}, \Psi = \Psi_L$

In presence of light, $\hat{H} = \hat{H}^{(0)} + \hat{H}^{(1)}, \Psi = C_L \Psi_L + C_m \Psi_m$

$$\hat{H} = -\mu x E_x$$

We are considering that it is an electrostatic interaction of electric field and electric dipole. We limit to only one axis.

$$E_x = 2E_x^0 \cos 2\pi \nu t, \nu = \text{frequency.}$$

$$[\hat{H}^{(0)} + \hat{H}^{(1)}] \Psi = -\frac{\hbar}{i} \frac{\partial}{\partial t} \Psi$$

$$[\hat{H}^{(0)} + \hat{H}^{(1)}] [C_L \Psi_L + C_m \Psi_m] = -\frac{\hbar}{i} \frac{\partial}{\partial t} [C_L \Psi_L + C_m \Psi_m]$$

$$\Rightarrow C_L \hat{H}^{(1)} \Psi_L + C_m \hat{H}^{(0)} \Psi_m + C_L \hat{H}^{(0)} \Psi_L + C_m \hat{H}^{(1)} \Psi_m = -\frac{\hbar}{i} \frac{\partial}{\partial t} [C_L \Psi_L + C_m \Psi_m]$$

$$= -\frac{\hbar}{i} \left[C_L \frac{\partial \Psi_L}{\partial t} + \frac{\partial C_L}{\partial t} \Psi_L + C_m \frac{\partial \Psi_m}{\partial t} + \frac{\partial C_m}{\partial t} \Psi_m \right]$$

$$C_L \hat{H}^{(1)} \Psi_L + C_m \hat{H}^{(0)} \Psi_m = -\frac{\hbar}{i} \left[\frac{\partial C_L}{\partial t} \Psi_L + \frac{\partial C_m}{\partial t} \Psi_m \right]$$

Now, multiply with Ψ_m^*

$$C_L \langle \Psi_L | \hat{H}^{(1)} | \Psi_L \rangle + C_m \langle \Psi_m | \hat{H}^{(0)} | \Psi_m \rangle = -\frac{\hbar}{i} \frac{\partial C_m}{\partial t}$$

$$\begin{aligned}\frac{\partial C_m}{\partial t} &= -\frac{i}{\hbar} \langle \psi_m | \hat{H}^{(1)} | \psi_k \rangle \\ &= -\frac{i}{\hbar} \langle \psi_m | \hat{H}^{(1)} | \psi_k \rangle \exp \left[\frac{i(E_m - E_k)t}{\hbar} \right] \exp \left[\frac{2\pi i(E_m - E_k)t}{h} \right]\end{aligned}$$

$$\begin{aligned}\hat{H} &= \hat{H}^{(0)} + \hat{H}^{(1)} \quad \hat{H}^{(1)} = -\mu_x E_x \quad E_x = 2E_x^0 \cos 2\pi \nu t \\ &= -\mu_x E_x^0 (\cos 2\pi \nu t)\end{aligned}$$

$$\begin{aligned}\hat{H}^{(1)} &= -\mu_x \cdot 2 \cdot E_x^0 \cdot \frac{1}{2} \left[\exp(2\pi i \nu t) + \exp(-2\pi i \nu t) \right] \\ &= -\mu_x E_x^0 \left[\exp(2\pi i \nu t) + \exp(-2\pi i \nu t) \right] \\ &= -\mu_x E_x^0 \left[\exp\left(\frac{2\pi i h \nu t}{h}\right) + \exp\left(-\frac{2\pi i h \nu t}{h}\right) \right]\end{aligned}$$

$$\begin{aligned}\frac{\partial C_m}{\partial t} &= i E_x^0 \langle \psi_m | \mu_x | \psi_k \rangle \left[\exp \left\{ \frac{2\pi i (E_m - E_k + h\nu)t}{h} \right\} \right. \\ &\quad \left. + \exp \left\{ \frac{2\pi i (E_m - E_k - h\nu)t}{h} \right\} \right]\end{aligned}$$

Molecular Spectroscopy

Jack D. Graybeal

$$\hat{H}^{(1)} = -\mu_x E_x$$

dipole moment
of the molecule

The amplitude of the electric field

$$E_x = \underbrace{2 E_x^0}_{\text{Maximum amplitude of the electric field}} \cos 2\pi \underbrace{\nu t}_{\text{frequency of radiation}}$$

The interaction between a dipole moment and electric field gives the perturbing potential.

$$\frac{dC_m}{dt} = + \frac{i E_x^0}{\hbar} \langle \psi_m | \mu_x | \psi_L \rangle \left[\exp \left\{ \frac{2\pi i}{\hbar} (E_m - E_L + h\nu) t \right\} + 2 \exp \left\{ \frac{2\pi i}{\hbar} (E_m - E_L - h\nu) t \right\} \right]$$

$$C_m = \frac{i}{\hbar} E_x^0 \langle \psi_m | \mu_x | \psi_L \rangle \left[\frac{1 - \exp \frac{2\pi i}{\hbar} (E_m - E_L + h\nu) t}{\frac{2\pi i}{\hbar} (E_m - E_L + h\nu)} + \frac{1 - \exp \frac{2\pi i}{\hbar} (E_m - E_L - h\nu) t}{\frac{2\pi i}{\hbar} (E_m - E_L - h\nu)} \right]$$

$$= E_x^0 \langle \psi_m | \mu_x | \psi_L \rangle \left[\frac{1 - \exp \left\{ \frac{2\pi i}{\hbar} (E_m - E_L + h\nu) t \right\}}{E_m - E_L + h\nu} + \frac{1 - \exp \left\{ \frac{2\pi i}{\hbar} (E_m - E_L - h\nu) t \right\}}{E_m - E_L - h\nu} \right]$$

For absorption,

$$E_m - E_L \cong h\nu$$

Emission,

$$E_m - E_L \cong -h\nu$$

So this term becomes significant

this term significant

$$C_m = E_x^0 \langle \psi_m | \mu_x | \psi_L \rangle \left[\frac{1 - \exp \left\{ \frac{2\pi i}{\hbar} (E_m - E_L - h\nu) t \right\}}{E_m - E_L - h\nu} \right]$$

The probability of a system existing in a given state having a wavefn ψ is proportional to $\psi^* \psi$.

$$\int \psi^* \psi dz = \int (C_L^* \psi_L^* + C_m^* \psi_m^*) (C_L \psi_L + C_m \psi_m) dz$$

$$= C_L^* C_L + C_m^* C_m$$

$$C_m = \frac{\pi t}{h} E_L^0 \langle \psi_m | \mu_x | \psi_L \rangle \left[\frac{1 - \exp \left\{ \frac{2\pi i}{h} (E_m - E_L - h\nu) t \right\}}{\pi (E_m - E_L - h\nu) \cdot t} \right]$$

$$= \frac{\pi t}{h} E_L^0 \langle \psi_m | \mu_x | \psi_L \rangle \left[\frac{1 - \exp i2A}{A} \right]$$

$$C_m^* C_m = \frac{\pi^2 t^2}{h^2} E_L^0 \langle \psi_m | \mu_x | \psi_L \rangle^2 \left[\frac{2 - e^{-i2A} - e^{i2A}}{A^2} \right]$$

Euler

$$e^{\pm i\gamma} = \cos \gamma \pm i \sin \gamma$$

$$= \frac{\pi^2 t^2}{h^2} E_L^0 \langle \psi_m | \mu_x | \psi_L \rangle^2 \left[\frac{2 - \cancel{\cos 2A} + i \cancel{\sin 2A} - \cancel{\cos 2A} - i \cancel{\sin 2A}}{A^2} \right]$$

$$= \frac{\pi^2 t^2}{h^2} E_L^0 \langle \psi_m | \mu_x | \psi_L \rangle^2 \left[\frac{2 - 2(1 - \sin^2 A)}{A^2} \right]$$

$$= \frac{\pi^2 t^2}{h^2} E_L^0 \langle \psi_m | \mu_x | \psi_L \rangle^2 \left[\frac{\sin^2 A}{A^2} \right]$$

Natural Line width

- 1) Collision broadening
- 2) Doppler broadening
- 3) Heisenberg uncertainty principle

$$\Delta E \times \Delta t \simeq \frac{\hbar}{2} = \frac{h}{2\pi} \quad 10^{-34} \text{ JS}$$

$$\delta \nu = \frac{\delta E}{h} \simeq \frac{h}{2\pi h \delta t} \simeq \frac{1}{2\pi \delta t}$$

excited electronic state lifetime 10^{-8} s

$$\delta \nu \simeq 10^8 \text{ Hz}$$

Usual radiation frequency $10^{14} - 10^{16}$

electronic spin state lifetime 10^{-7} s

$$\delta \nu \simeq 10^7 \text{ Hz}$$

Usual radiation frequency $10^8 - 10^9 \text{ Hz}$

Intensity of spectral lines

- 1) Transition probability
- 2) Population of states
- 3) Path length of sample

$$\frac{I}{I_0} = e^{-\epsilon \cdot c \cdot l}$$

Beer-Lambert Law

Electronic spectroscopy

How do we define an electronic state?

Born-Oppenheimer Approximation

$$\Psi = \Psi_{es} \Psi_v \Psi_r$$

$$E = E_{es} + E_v + E_r$$

— S_1 — T_1

— S_0

$\alpha(1)\alpha(2)$
 $\beta(1)\beta(2)$

HOMO
↑
Singlet

LUMO
↓
Singlet

↑
Triplet

$\alpha(1)\beta(2) + \beta(1)\alpha(2)$ } symmetric with exchange.
 $\alpha(1)\beta(2) - \beta(1)\alpha(2)$

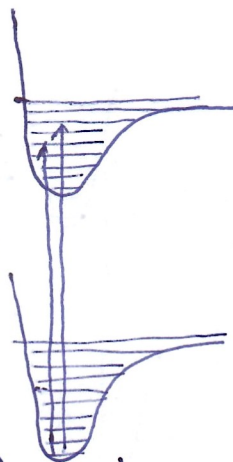
Triplet state is a combination of 3 wavefns.
Singlet " " a single wavefn.

(can we experimentally see?
Yes, by applying magnetic field)

$$\Psi = \Psi_{es} \Psi_v$$

$$E = E_{es} + E_v$$

The resulting state is called a Franck-Condon state, and the transition involves a vertical transition.



Franck-Condon Principle

* James Franck 1926
* Edward Uhler Condon
Classically the Franck-Condon principle is the approximation that an electronic transition is most likely to occur without

changes in the positions of the nuclei

$$TMI = \langle \Psi_{es}^{(2)} \Psi_v^{(2)} | (\mu_N + \mu_e) | \Psi_{es}^{(1)} \Psi_v^{(1)} \rangle$$

$$= \langle \Psi_{es}^{(2)} \Psi_v^{(2)} | \mu_N | \Psi_{es}^{(1)} \Psi_v^{(1)} \rangle + \langle \Psi_{es}^{(2)} \Psi_v^{(2)} | \mu_e | \Psi_{es}^{(1)} \Psi_v^{(1)} \rangle$$

$$= \langle \Psi_{es}^{(2)} | \Psi_{es}^{(1)} \rangle \langle \Psi_v^{(2)} | \mu_N | \Psi_v^{(1)} \rangle + \langle \Psi_{es}^{(2)} | \mu_e | \Psi_{es}^{(1)} \rangle \langle \Psi_v^{(2)} | \Psi_v^{(1)} \rangle$$

↑ electronic wavefn of same different

$$TM1 = \langle \psi_e^{(2)} | \psi_e^{(1)} \rangle$$

Franck-Condon factor

If good overlap, then

$$\psi_{es} = \psi_e \cdot \psi_s$$

$$\langle \psi_e^{(2)} | \mu_e | \psi_e^{(1)} \rangle \langle \psi_s^{(2)} | \psi_s^{(1)} \rangle$$

↓ symmetry
of electronic
wave fns.

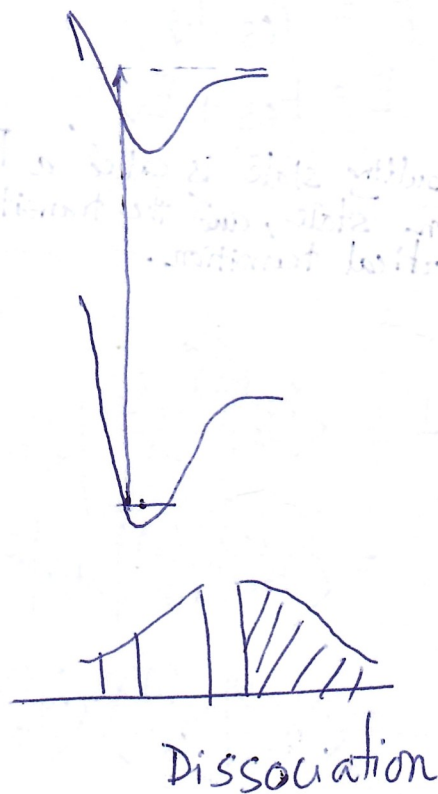
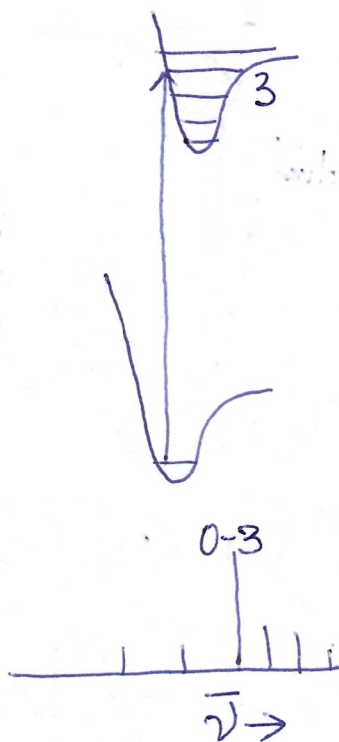
Orbital forbidden/allowed

↓ $\Delta S = 0, \Delta m_s = 0$

Spin forbidden/allowed.

* Orbital forbidden transitions can be vibronically allowed.

* Spin forbidden transitions can become allowed.
How? Spin-orbit splitting.



The quantum mechanical formulation of this principle is that the intensity

Electronic Spectroscopy: Franck-Condon Factor

Born-Oppenheimer $\Psi = \Psi_{es} \Psi_v \Psi_r$

$$\downarrow$$

$$\Psi_e \cdot \Psi_s$$

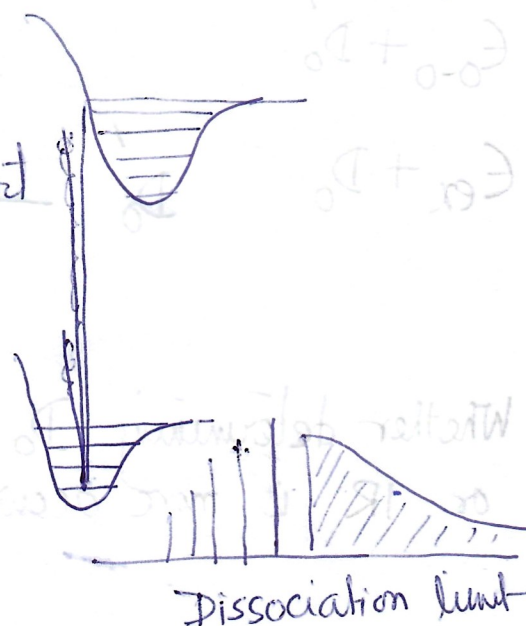
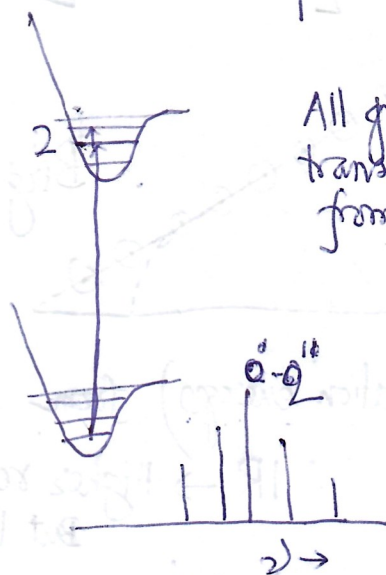
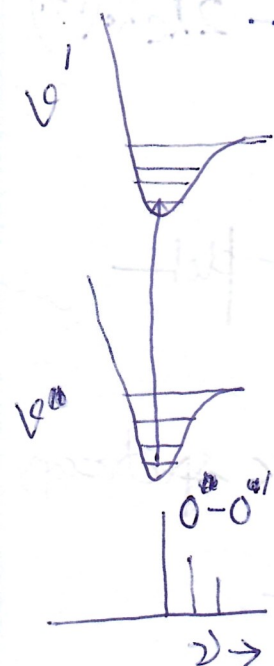
$$TMI = \langle \Psi_v^{(2)} | \Psi_v^{(1)} \rangle \langle \Psi_{es}^{(2)} | \mu | \Psi_{es}^{(1)} \rangle \langle \Psi_s^{(2)} | \Psi_s^{(1)} \rangle$$

Franck-Condon factor χ (Chi) Orbital Selection Spin selection

Spin & orbitally allowed	3 5 10 to 10
Spin allowed but orbitally forbidden	10 ⁰ to 10 ³
Spin forbidden but orbitally allowed	10 ⁻⁵ to 10 ⁰

Franck-Condon Principle

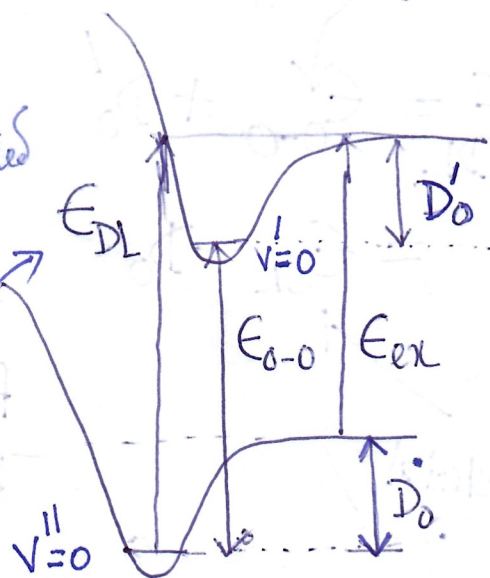
All ground transitions start from $v=0$



D_0' = Dissociation energy for excited state

E_{0-0} Band origin

Dissociation limit



E represents energy

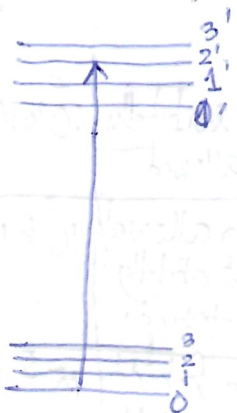
$$E_{DL} = E_{0-0} + D_0'$$

If dissociation limit is very well-defined and E_{0-0} is clear we can determine D_0' . But usually it is not the case.

Thermochemically we can measure directly, the excited state energy

$$E_{DL} = E_{ex} + D_0$$

If we don't have dissociation limit in the spectrum, then how can we determine



Born-Oppenheimer approximation

$$\epsilon = \epsilon_{es} + \epsilon_v$$

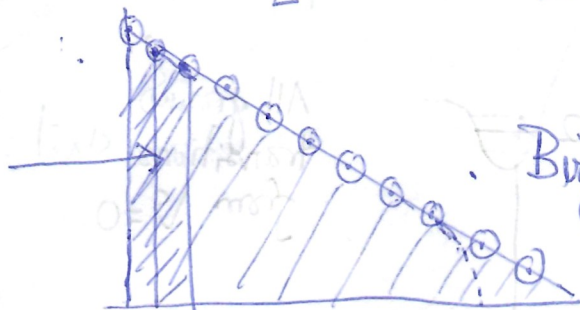
$$\Delta \epsilon = \epsilon_{00} + \Delta \epsilon_v$$

$$\epsilon_v = \left[\left(v + \frac{1}{2} \right) \bar{\nu}_v \chi_e - \left(v + \frac{1}{2} \right)^2 \bar{\nu}_v \bar{\chi}_e \right]$$

$0 \rightarrow v'$
 $0 \rightarrow v'+1$

Two transitions.

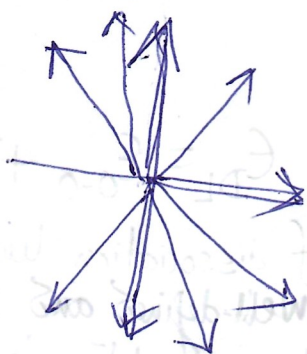
$$\Delta \epsilon = \bar{\nu}_v \left[1 - 2\chi_e(v'+1) \right] = \bar{\nu}_v - 2\chi_e \bar{\nu}_v (v'+1)$$



Birge-Sponer plot

Whether determining D_0 (dissociation energy) from electronic spectroscopy or IR is more accurate?

IR $\rightarrow$ Higher resolution
But less points



Band origin

$$\epsilon_{DL} = \epsilon_{0-0} + D_0'$$

$$= \epsilon_{ex} + D_0$$

Can be determined from thermochemical

R.T. Birge

Raymond Thayer Birge

Hertha Sponer